# INCORPORATING CCSDS TELEMETRY STANDARDS AND PHILOSOPHY ON CASSINI

John C.Day and Anne B.Elson

Jet Propulsion Laboratory, California Institute of "l'ethnology
Pasadena, «California

### **Abstract**

The Cassini project at the Jet Propulsion Laboratory (IPL) is implementing a spacecraft telemetry system based on the Consultative Committee for Space Data Systems (CCSDS) packet telemetry standards. Resolving the CCSDS concepts with a Ground Data System designed to handle time-division-multiplexed telemetry and also handling constraints unique to a deep-space planetary spacecraft (such as fixed downlink opportunities, small downlink rates and requirements for on-board data storage) have resulted in spacecraft and ground system design challenges. Solving these design challenges involved adapting and extending the CCSDS telemetry standards as well as changes to the spacecraft and ground system designs. The resulting spacecraft/ground system design is an example of how new ideas and philosophies can be incorporated into existing systems and design approaches without requiring significant rework. In addition, it shows that the CCSDS telemetry standards can be successfully applied to deep-space planetary spacecraft.

#### 1.0 Introduction

In order to provide context, this section describes the Cassini Mission, Cassini Spacecraft, and Cassini Ground System.

### 1.1 Cassini Mission

The goal of the Cassini mission is to provide in-depth exploration of the Saturnian system. This includes the planet itself, the rings and magnetosphere, the moon Titan, and eight icy satellites. Launch of the Cassini spacecraft is set for October 1997. The trajectory to Saturn will take 6.7 years and require flybys of Venus and Earth for gravity assists. The trajectory also includes a Jupiter flyby. Arriving at Saturn in June 2004, the spacecraft will perform a Saturn Orbit Insertion (S01) maneuver and will orbit the planet for four years, until June 2008. This will provide a 60 orbit tour that includes 33 Titan flybys. During the first or second flyby, the Huygens probe (provided by ESA) will be released to study the atmosphere of the satellite. On some of the remaining Titan flybys, a synthetic aperture radar (SAR) will be used to penetrate the obscuring atmosphere of Titan to take SAR images of the moon's surface.

### 1.2 Cassini Spacecraft

The Cassini spacecraft is approximately 6.60 meters (m) long, 4.0 m in diameter, and has a mass allocation of 5,655 kilograms (kg) (2,523 for the spacecraft, and 3,132 for fuel). The spacecraft is subdivided into two major components - the orbiter and the Huygens probe. The orbiter delivers the Huygens probe to Saturn's moon Titan, then continues with the reconnaissance of Saturnian system. Cassini is somewhat unique for a planetary spacecraft in that it has no articulated platforms. The orbiter must orient itself in order to point any directional device, e.g. cameras or antennas.

The spacecraft contains six major engineering subsystems:

- 1) Radio Frequency Subsystem (RFS)
  - uplink command reception
  - telemetry transmission

# 2) Command and Data Subsystem (CDS)

- uplink command processing
- spacecraft intercommunication
- telemetry packet collection
- on-board telemetry storage
- telemetry frame creation
- 3) Attitude and Articulation Control Subsystem (AACS)
  - attitude determination
  - attitude control
  - thrust vector control
  - main engine control
- 4) Propulsion Module Subsystem (PMS)
  - propellant tanks
  - thrusters
  - main engines
- 5) Power and Pyrotechnics Subsystem (PPS)
  - power supply, conditioning and control
  - pyrotechnic firing circuitry
- 6) Probe Support Avionics (PSA)
  - interface to Huygens Probe
  - probe commanding
  - probe mission telemetry

'] 'he Cassini orbiter's science payload is a set of twelve science instruments (the Huygens probe also has its own science payload, which will not be discussed here, as collection of this science data is managed by the probe). The twelve instruments are made up of four optical remote sensing science instruments, six fields, particles and waves science instruments, and two microwave remote sensing science instruments.

The spacecraft is relatively quiescent during its cruise to Saturn and is in communication with Earth usually only once per week. Once in Saturn orbit, the orbiter will be in daily contact with the ground. While in orbit around Saturn, the orbiter will collect science data for sixteen hours and then will point the high gain antenna at the earth to

downlink the data for eight hours. There are eight downlink rates used for transmitting the science data during the tour phase. These rates vary from 14,220 to 165,900 bps. In order to accommodate data collection while not pointing at Earth, the orbiter has the capability to store up to four gigabits ofdata on a solid-state recorder (SSR).

# 1.3 Cassini Ground System (Telemetry)

The telemetry processing portion of the Cassini Ground System is based upon a system developed for multi-mission operations. I-his system was used with the Magellan, Mars Observer and Galileo spacecraft. Adaptations to this system to support Cassini were planned, but no architectural changes were to be made.

The multi-mission system was designed to support time-division-multiplexed (TDM) telemetry structures used on previous spacecraft programs such as Voyager, Galileo and Magellan. TDM telemetry is characterized by fixed-size telemetry frames, which contain measurements in a fixed order, with only very specific and limited variations (e.g., subcommutation). These frames are transmitted to the ground at only a small number of quantit ized downlink rates. Each frame contains a synchronization pattern which provides the ground system a means for determining frame boundaries.

The resulting multi-mission system, used as the basis for the Cassinitelemetry processing system, placed strict limits on the flexibility of telemetry production on the spaceci aft. For example, these strictures prevented the use of variable size transfer frames, placed restrictions on telemetry packet size variability and measurement placement, did not allow time-tagging below the packet level, and did not easily handle multiple measurements in a single packet (supercommutation). Incorporating CCSIDS telemetry design concepts, which encourage telemetry systems to be extremely flexible, in this environment presented many challenges.

# 2.0 Process for Incorporating CCSDS Standards

The incorporation of the CCSDS standards and philosophy into spacecraft telemetry systems at JPL is a continuing process which has progressed through many stages. The Mars Observer (MO) spacecraft was the first JPL spacecraft to use the CCSDS telemetry standards. Cassini is the second application, and while the Cassini telemetry design is constrained in many similar ways as the MO design, it also moves closer to implementing the CCSDS telemetry philosophy.

## 2.1Difficulties with CCSDS Philosophy

The Cassini project decided to use the CCSDS standards for both the command and telemetry functions. However, full incorporation of the telemetry standards and philosophyheld some challenges due to limitations of the Cassini mission, necessary characteristics of deep-space spacecraft, and differences between the CCSDS design concepts and the JPL culture.

The CCSDS telemetry standards provide a generic protocol for transferring observational data from the instrument to the end-user. This protocol was designed to facilitate "the acquisition and transmission of instrument data at a rate appropriate to the phenomenon being observed" <sup>1</sup>. This implies a design concept of demand-driven telemetry, where the tclcmctry producers determine the rate and size of the telemetry packets they produce, instead of a centralized packet collector. In addition, the CCSDS standards form a layeredtelemetry architecture, in which each layer depends on a set of well-defined functions provided by the neighboring layers. This design concept focuses on the telemetry functions to be performed, rather than the form of the telemetry system, and allows the physical design to be decoupled from the telemetry collection process.

The nature of the Cassini mission involves constraints which do not apply to other spacecraft, in particular the earth-orbiting spacecraft for which the CCSDS standards were primarily designed. These constraints include a

limited downlink rate and constrained transmission opportunities, which in turn affect the design of the on-board data storage system.

The trajectory for the Cassini mission requires telemetry links at distances up to 10 AU over three antennas (1 high-gain, 2 lowgain), many times with a constrained attitude due to thermal limitations. The downlink rates for the first two years of the mission (with exception of onc month-long period) range from 948 bits/see down to 20 bits/see. The limited downlink rate is an implicit requirement for efficient collection and transmission of the spaceci aft telemetry. Because the CCSDS standards require considerable overhead (96 bits per telemetry packet, plus frame overhead), the average size of a Cassini telemetry packet is approximately 8000 bits. The large size results in a header overhead which is approximately 15% (considered to be areasonable amount). However, at the low downlink rates during the early part of the mission, large packet sizes and large flame size (the Cassini frame size is fixed at 10112. bits) create problems with frame synchronization and measurement time-tagging. The large transfer frame size causes long intervals between reception of sync signals at low downlink rates, which increases the likelihood of a ground glitch or transmission channel noise causing an error which causes loss of the frame. Large packet sizes result in long periods between measurement updates (since supercommutated measurements cannot be uniquely time-tagged, cach packet contains only one instance of a measurement).

The constrained downlink opportunities come from the use of the Deep Space Network (DSN), a necessary resource for deep-space missions which is shared with many other spacecraft. The availability of the DSN receiving stations drives the spacecraft data storage design. A spacecraft design which produces a regular and predictable amount of telemetry is required as the scheduling of the DSN stations is done months in advance and only in emergency situations can a station be requested "on-the-fly" (as opposed to most earth-orbiting spacecraft which have continuous covers ge). Therefore the maximum volume of

data collected between downlink passes must be known in advance to avoid losing data due to oversubscription of the available storage volume. It is difficult to develop an efficient telemetry system with this constraint, given that science instruments do not always produce predictable amounts of data, due to the type of data being taken or variable-ratio compression algorithms. The collection and playback of telemetry must be balanced with the available downlink opportunities.

IPL has a history of producing extremely reliable spacecraft to perform the deep-space science missions traditionally associated with IPL. A large part of ensuring the reliability of IPL spacecraft is in the testing and verification done pre-launch. IPL spacecraft are designed to perform their functions in a simple, reliable and deterministic manner. The TDM telemetry system used on previous JPL projects in completely deterministic, and can be verified in a straightforward manner. In contrast, the systems envisioned by the CCSDS, which use concepts such as demand-driven telemetry, arc more difficult (and time-consuming) to test rigorously. Duc to this issue, there is a hesitancy at JPL to develop telemetry systems which fully utilize the potential of packet telemetry systems. For example, the telemetry system on Mars Observer spacecraft, which implemented the CCSDS telemetry standards, required all telemetry packet producers to use the same packet size at all times. The experience of building the MO design enabled the Cassini telemetry system to take another step toward dctnand-driven telemetry, but it is still limited in its implementation of the design concepts endorsed by the CCSDS.

### 2.2 Architectural Decisions

In the Cassini pre-project phase, a team was formed to discuss how much of the CCSDS telemetry standards would be incorporated in the Cassini design. The decisions of this team, whose members included developers of the MO telemetry system, formed the basis of the Cassini spacecraft telemetry system. The protocol of the CCSDS telemetry standards was accepted, but the concept of a demand-driven telemetry

system was not. One of the most significant architectural decisions was to develop a schedule-driven packet collection scheme, rather than a demand-driven. Under this scheme, a spacecraft telemetry mode (STM) table would be defined which would control the packet collection, packet routing and associated downl ink rate. Within a S'I'M, packet size and rate from each source is fixed, but are allowed to be different in other STMs. A set of STMs would be developed to provide for the varied telemetry needs over the course of the mission. Changes to any portion of the spacecraft telemetry mode specification (collection, routing or downlink rate) required the definition of a new spacecraft telemetry mode. "I-his approach, although it allows a packet telemetry system to be used on a planetary spacecraft, still severely constrains the spacecraft's ability to allow data producers to independently determine and vary both the rate and amount of telemetry data they produce over time.

### 2.3 System Design

With this architecture in place, the designers of the spacecraft information system began the process of developing in more detail the spacecraft telemetry system. It soon became evident that additional capabilities, which expanded on the architectural design without violating the architectural principles, testability or ground system capabilities, would be required in order to have a viable telemetry system design. Many of these extensions incorporated elements of the CCSDS telemetry philosophy.

An important capability is the ability of a telemetry producer to vary the contents (format) of their telemetry packets. Although the CDS collects telemetry from each producer at a fixed rate and fixed packet size (within a STM), if the size and rate remain the same, the instrument or subsystem may change the contents of the packets produced based on internal mode changes. "I-his allows some telemetry variability within a given STM. The packet variations are identified by a different application process identifier (APID) in the packet header. However, since only a limited number of bits is available for differentiating

packet types, each subsystem is limited to at most 64 packet variations.

Another capability available for packet producers to vary the contents of their packets is through the usc of mini-packets. Mini-packets arc an extension of the CCSDS packet telemetry standards which allow for greater flexibility in placement of measurements within a packet. Although changing the packet type provides some variability, a significant restriction is that within a packet type, measurement positions are fixed (due to constraints of the ground system). "l'his, in combination with the limited number of packet types allowed per subsystem, constrains the variability in a given subsystem's or instrument's telemetry. 'l'his was especially restrictive to the AACS team, since the AACS must provide accurate and regular telemetry for correlation of science data and analysis of control system performance (e.g., attitude and rate data). Mini-packets arc small collections of related measurements which have all been sampled within a certain, small time-frame. Mini-packets were first used on Mars Observer. but their implementation was constrained limited to a fixed number of orientations. Cassini has extended the mini-packet concept to allow placement of mini-packets anywhere within the packet, enabling much more flexible and event-driven telemetry<sup>2</sup>. In addition, Cassini mini-packets contain small time-tags, which when combined with the packet time-tag, provide fine timing resolution within the packet (i.e., beyond the accuracy provided by the packet time-tag). As a result, many instances of the same mini-packet can be collected and placed in the same packet, allowing measurement sampling at a rate greater than the packet collection rate, without the difficulties of supercommutation. The Cassini implementation of mini-packets required some modifications to the ground system software, but due to the support of ground system developers during the process, a design which provided the required capabilities to the spacecraft was developed with minimal cost and impacts.

In order to provide packet producers with the capability to adjust the rate as well as

the contents of telemetry packets, a concept known as "zero-lct~gtb packets" (ZLPs) was introduced. ZLPs are packets produced when an instrument has not collected enough information to provide a packet at the scheduled packet pickup. ZLPs are identified by setting the length field in the packet header to an invalid (too small) value, but the actual size of the packet is the size defined for that producer in the current STM.ZLPs are collected by the CDS, but are not stored or downlinked. This allows instruments to use only as much of their packet collection rate allocation in a given STM as needed. The packet collection rate allocation is usually set at the maximum data rate the instrument is allowed to produce in that S-l<M, so the instruments can vary the effective output data rate without affecting the spacecraft telemetry mode. This allows instruments with variable-ratio compression algorithms to uscless downlink bandwidth and avoid storing packets which contain no data on the SSR. This capability, which gives the packet producers limited control over the volume of data produced, is a significant step toward allowing demand-driven telemetry.

## 3.0 Resulting Capabilities

As a result of the implementation of the Cassini telemetry system, many of the capabilities envisioned by the CCSDS have been made possible within the mission and architectural constraints. Additionally, there have been some resulting benefits which increase the quality of the spacecraft telemetry.

While the Cassini spacecraft telemetry system began as a very limited implementation of the CCSDS standards and philosophies, the additional capabilities which were incorporated show an evolution toward the CCSDS design concepts. The STMS used by Cassini limit the number of telemetry producers at a given time and also dictate the amount of data each producer is allowed to produce. However, the zero-length packet concept allows a telemetry producer to vary the actual data volume within the limit set by the STM. In addition, changing the packet type or the use of mini-packets gives telemetry producers a great deal of control over

the telemetry produced. When these concepts are used effectively by the telemetry producer, the very efficient use of the limited telemetry resources can be accomplished.

In addition, the mini-packet concept provided a secondary benefit to the end-users of the spacecraft telemetry - greater measurement intercorrelation. One of the results of the large size of the Cassini telemetry packets is coarse indication of measurement sampling time. Evenat high downlink rates, several seconds may pass between the collection of successive packets from an engineering subsystem. Because measurements arc time-tagged with the packet time, it is difficult to determine exactly when a measurement changes values or when measurements within a packet were sampled with respect to one another. However, minipackets can be used to sample measurements faster than once per packet. This allows for much better resolution of important measurements and better intercorrelation between measurements because the mini-packet time-tag applies to a smaller number of measurements.

### 4.0 Conclusions

"I-he CCSDS telemetry standards can be successfully applied to deep-space planetary missions, even in an environment of rigorous testing, design conservatism and biased culture. The Cassinite lemetry system is an example of this, as it incorporates the CCSDS protocols and even with severe architectural constraints, is a flexible telemetry system not unlike the systems envisioned by the CCSDS designers. "l"his was accomplished by developing solutions which allowed variability in the spacecraft telemetry while remaining within the architectural! tenets of the design and the constraints of the mission. The Cassini design solutions, and the solutions used on the Mars Observer spacecraft, reflect an evolution in IPL spacecraft telemetry systems away from the traditional time-division-multiplexed systems toward systems which completely embrace the CCSDS standards and philosophy.

## **5.0** Acknowledgments

The work described in this paper was carried out at the JetPropulsion Laboratory, California Institute of-l'ethnology, under a contract with the National Aeronautics and Space Administration. The implementation of the Cassini spacecraft telemetry system is based upon early work in the spacecraft information system done by Sandy Krasner.

### 6.0 References

- 1 "Summary of Concept and Rationale", Report CCSDS 100.0-C-1, Issue 1, Green Book, Consultative Committee for Space Data Systems, December 1987.
- 2 J. C.Day, "Telemetry Formats and Measurements", JPLPD 699-205-3-281, May 1994 (JPL Internal 1 Document)

# INCORPORATING CCSDS TELEMETRY STANDARDS AND PHILOSOPHY ON CASSINI

John C. Day and Anne B. Elson Jet Propulsion Laboratory, California institute of Technology Pasadena, California

#### Abstract

The Cassini project at the Jet Propulsion I Laboratory (JPL) is implementing a spacecraft telemetry system based on the Consultative Committee for Space Data Systems (CCSDS) packet telemetry standards. Resolving the CCSDS concepts with a Ground Data System designed to handle time-division-multiplexed telemetry and also handling constraints unique to a deep-space planetary spacecraft (such as fixed downlink opportunities, small downlink rates and requirements for on-board data storage) have resulted in spacecraft and ground system design challenges. Solving these design challenges involved adapting and extending the CCSDS telemetry standards as well as changes to the spacecraft and ground system designs. The resulting spacecraft/ground system design is an example of how new ideas and philosophies can be incorporated into existing systems and design approaches without requiring significant rework, In addition, it shows that the CCSDS telemetry standards can be successfully applied to deep-space planetary spacecraft.

### 1.0 introduction

In order to provide context, this section describes the Cassini Mission, Cassini Spacecraft, and Cassini Ground System.

## 1.1 Cassini Mission

The goal of the Cassini mission is to provide in-depth exploration of the Saturnian system. This includes the planet itself, the rings and magnetosphere, the moon Titan, and eight icy satellites. Launch of the Cassini spacecraft is set for October 1997. The trajectory to Saturn will take 6.7 years and require flybys of Venus and Eat th for gravity assists. The trajectory also includes a Jupiter flyby. Arriving at Saturn in June 2004, the spacecraft will perform a Saturn Orbit Insertion (S01) maneuver and will orbit the planet for four years, until June 2008. This will provide a 60 orbit tour that includes 33 Titan flybys. During the first or second flyby, the Huygens probe (provided by ESA) will be released to study the atmosphere of the satellite. On some of the remaining Titan flybys, a synthetic aperture radar (SAR) will be used to penetrate the obscuring atmosphere of Titan to take SAR images of the moon's surface.

# 1.2 Cassini Spacecraft

The Cassini spacecraft is approximately 6.60 meters (m) long, 4.0 m in diameter, and has a mass allocation of 5,655 kilograms (kg) (2,523 for the spacecraft, and 3,132 for fuel). "I"he spacecraft is subdivided into two major components - the orbiter and the Huygens probe. The orbiter delivers the Huygens probe to Saturn's moon '1 "itan, then continues with the reconnaissance of Saturnian system. Cassini is somewhat unique for a planetary spacecraft in that it has no articulated platforms. The orbiter must orient itself in order to point any directional device, e.g. cameras or antennas.

The spacecraft contains six major engineering subsystems:

- 1) Radio Frequency Subsystem (RFS)
  - uplink command reception
  - telemetry transmission

1

- 2) Command and Data Subsystem (CDS)
  - uplink command processing
  - spacecraft intercommunication
  - telemetry packet collection
  - on-board telemetry storage
  - telemetry frame creation
- 3) Attitude and Articulation Control Subsystem (AACS)
  - attitude determination
  - attitude control
  - thrust vector control
  - main engine control
- 4) Propulsion Module Subsystem (PMS)
  - propellant tanks
  - thrusters
  - main engines
- 5) Power and Pyrotechnics Subsystem (PPS)
  - power supply, conditioning and control
  - pyrotechnic firing circuitry
- 6) Probe Support Avionics (PSA)
  - interface to H u ygens Probe
  - probe commanding
  - probe mission telemetry

The Cassini orbiter's science payload is a set of twelve science instruments (the Huygens probe also has its own science payload, which will not be discussed here, as collection of this science data is managed by the probe). The twelve instruments are made up of four optical remote sensing science instruments, six fields, particles and waves science instruments, and two microwave remote sensing science instruments.

The spacecraft is relatively quiescent during its cruise to Saturn and is in communication with Earth usually only once per week. Once in Saturn orbit, the orbiter will be in daily contact with the ground. While in orbit around Saturn, the orbiter will collect science data for sixteen hours and then will point the high gain antenna at the earth to

downlink the data for eight hours. There are eight downlink rates used for transmitting the science data during the tour phase. These rates vary from 14,220 to 165,900 bps. In order to accommodate data collection while not pointing at Earth, the orbiter has the capability to store up to four gigabits of data on a solid-state recorder (SSR).

# 1.3 Cassini Ground System (Telemetry)

The telemetry processing portion of the Cassini Ground System is based upon a system developed for multi-mission operations. This system was used with the Magellan, Mars Observer and Galileo spacecraft. Adaptations to this system to support Cassini were planned, but no architectural changes were to be made.

The multi-mission system was designed to support time-division-multiplexed (TDM) telemetry structures used on previous spacecraft programs such as Voyager, Galileo and Magellan. TDM telemetry is characterized by fixed-size telemetry frames, which contain measurements in a fixed order, with only very specific and limited variations (e.g., subcommutation). These frames are transmitted to the ground at only a small number of quantitized downlink rates. Each frame contains a synchronization pattern which provides the ground system a means for determining frame boundaries.

The resulting multi-mission system, used as the basis for the Cassini telemetry processing system, placed strict limits on the flexibility of telemetry production on the spacecraft. For example, these strictures prevented the use of variable size transfer frames, placed restrictions on telemetry packet size variability and measurement placement, did not allow time-tagging below the packet level, and did not easily handle multiple measurements in a single packet (supercommutation). Incorporating CCSDS telemetry design concepts, which encourage telemetry systems to be extremely flexible, in this environment presented many challenges.

# 2.0 Process for Incorporating CCSDS Standards

The incorporation of the CCSDS standards and philosophy into spacecraft telemetry systems at JPL is a continuing process which has progressed through many stages. The Mars Observer (MO) spacecraft was the first JPL spacecraft to use the CCSDS telemetry standards. Cassini is the second application, and while the Cassini telemetry design is constrained in many similar ways as the MO design, it also moves closer to implementing the CCSDS telemetry philosophy.

## 2.1 Difficulties with CCSDS Philosophy

"I'he Cassini project decided to use the CCSDS standards for both the command and telemetry functions. However, full incorporation of the telemetry standards and philosophy held some challenges due to limitations of the Cassini mission, necessary characteristics of deep-space spacecraft, and differences between the CCSDS design concepts and the JPL culture.

The CCSDS telemetry standards provide a generic protocol for transferring observational data from the instrument to the end-user. This protocol was designed to facilitate "the acquisition and transmission of instrument data at a rate appropriate to the phenomenon being observed" <sup>1</sup>. 'l'his implies a design concept of demand-driven telemetry, where the telemetry producers determine the rate and size of the telemetry packets they produce, instead of a centralized packet collector. In addition, the CCSDS standards form a layered telemetry architecture, in which each layer depends on a set of well-defined functions provided by the neighboring layers. This design concept focuses on the telemetry functions to be performed, rather than the form of the telemetry system, and allows the physical design to be decoupled from the telemetry collection process.

"I'he nature of the Cassini mission involves constraints which do not apply to other spacecraft, in particular the earth-orbiting spacecraft for which the CCSDS standards were primarily designed. These constraints include a limited downlink rate and constrained transmission opportunities, which in turn affect the design of the on-hoard data storage system.

The trajectory for the Cassini mission requires telemetry links at distances up to 10 AU over three antennas (1 high-gain, 2 lowgain), many times with a constrained attitude due to thermal limitations. The downlink rates for the first two years of the mission (with exception of one month-long period) range from 948 bits/see down to 20 bits/see. The limited downlink rate is an implicit requirement for efficient collection and transmission of the spacecraft telemetry. Because the CCSDS standards require considerable overhead (96 bits per telemetry packet, plus frame overhead), the average size of a Cassini telemetry packet is approximately 8000 bits. The large size results in a header overhead which is approximately 15?40 (considered to be areasonable amount). However, at the low downlink rates during the early part of the mission, large packet sizes and large frame size (the Cassini frame size is fixed at 10112 bits) create problems with frame synchronization and measurement time-tagging. The large transfer frame size causes long intervals between reception of sync signals at low downlink rates, which increases the likelihood of a ground glitch or transmission channel noise causing an error which causes loss of the frame. Large packet sizes result in long periods between measurement updates (since supercommutated measurements cannot be uniquely time-tagged, each packet contains only one instance of a measurement).

The constrained downlink opportunities come fi om the use of the Dccp Space Network (DSN), a necessary resource for deep-space missions which is shared with many other spacecraft. The availability of the DSN receiving stations drives the spacecraft data storage design. A spacecraft design which produces a regular and predictable amount of telemetry is required as the scheduling of the DSN stations is done months in advance and only in emergency situations can a station be requested "on-the-fly" (as opposed to most earth-orbiting spacecraft which have continuous coverage). Therefore the maximum volume of

data collected between downlink passes must be known in advance to avoid losing data due to oversubscription of the available storage volume. It is difficult to develop an efficient telemetry system with this constraint, given that science instruments do not always produce predictable amounts of data, due to the type of data being taken or variable-ratio compression algorithms. The collection and playback of telemetry must be balanced with the available downlink opportunities.

IPI, has a history of producing extremely reliable spacecraft to perform the deep-space science missions traditionally associated with JPL. A large part of ensuring the reliability of IPL spacecraft is in the testing and verification done prc-launch. IPL spacecraft are designed to perform their functions in a simple, reliable and deterministic manner. The TDM telemetry system used on previous JPL projects in completely deterministic, and can be verified in a straightforward manner. In contrast, the systems envisioned by the CCSDS, which use concepts such as demand-driven telemetry, are more difficult (and time-consuming) to test rigorously. Duc to this issue, there is a hesitancy at IPL to develop telemetry systems which fully utilize the potential of packet telemetry systems. For example, the telemetry system on Mars Observer spacecraft, which implemented the CCSDS telemetry standards, required alltclcmetry packet producers to use the same packet size at all times. The experience of building the MO design enabled the Cassini telemetry system to take another step toward demand-driven telemetry, but it is still limited in its implementation of the design concepts endorsed by the CCSDS.

### 2.2 Architectural Decisions

in the Cassini pre-project phase, a team was formed to discuss how much of the CCSDS telemetry standards would be incorporated in the Cassini design. The decisions of this team, whose members included developers of the MO telemetry system, formed the basis of the Cassini spacecraft telemetry system. The protocol of the CCSDS telemetry standards was accepted, but the concept of a demand-driven telemetry

system was not. One of the most significant architectural decisions was to develop a schedule-driven packet collection scheme, rather than a demand-driven. Under this scheme, a spacecraft telemetry mode (STM) table would be defined which would control the packet collection, packet routing and associated downlink rate. Within a STM, packet size and rate from each source is fixed, but arc allowed to be different in other STMs. A set of STMs would be developed to provide for the varied telemetry needs over the course of the mission. Changes to any portion of the spacecraft telemetry mode specification (collection, routing or downlink rate) required the definition of a new spacecraft telemetry mode. This approach, although it allows a packet telemetry system to be used on a planetary spacecraft, still severely constrains the spacecraft's ability to allow data producers to independently determine and vary both tile rate and amount of telemetry data they produce over time.

# 2.3 System Design

With this architecture in place, the designers of the spacecraft information system began the process of developing in more detail the spacecraft telemetry system. It soon became evident that additional capabilities, which expanded on the architectural design without violating the architectural principles, testability or ground system capabilities, would be required in order to have a viable telemetry system design. Many of these extensions incorporated elements of the CCSDS telemetry philosophy.

An important capability is the ability of a telemetry producer to vary the contents (format) of their telemetry packets. Although the CDS collects telemetry from each producer at a fixed rate and fixed packet size (within a STM), if the size and rate remain the same, the instrument or subsystem may change the contents of the packets produced based on internal mode changes. "I'his allows some telemetry variability within a given STM. The packet variations are identified by a different application process identifier (APID) in the packet header. However, since only a limited number of bits is available for differentiating

packet types, each subsystem is limited to at most 64 packet variations.

Another capability available for packet producers to vary the contents of their packets is through the use ofmini-packets. Mini-packets are an extension of the CCSDS packet telemetry standards which allow for greater flexibility in placement of measurements within a packet. Although changing the packet type provides some variability, a significant restriction is that within a packet type, measurement positions are fixed (due to constraints of the ground system). This, in combination with the limited number of packet types allowed per subsystem, constrains the variability in a given subsystem's or instrument's telemetry. This was especially restrictive to the AACS team, since the AACS must provide accurate and regular telemetry for correlation of science data and analysis of control system performance (e.g., attitude and rate data). Mini-packets arc small collections of related measurements which have all been sampled within a certain, small time-frame. Mini-packets were first used on Mars Observer, but their implementation was constrained limited to a fixed number of orientations. Cassini has extended the mini-packet concept to allow placement of mini-packets anywhere within the packet, enabling much more flexible and even t-d riven telemetry<sup>2</sup>. In addition, Cassini mini-packets contain small time-tags. which when combined with the packet time-tag, provide fine timing resolution within the packet (i.e., beyond the accuracy provided by the packet time-tag). As a result, many instances of the same mini-packet can be collected and placed in the same packet, allowing measurement sampling at a rate greater than the packet collection rate, without the difficulties of supercommutation. The Cassini implementation of mini-packets required some modifications to the ground system software, but due to the support of ground system developers during the process, a design which provided the required capabilities m the spacecraft was developed with minimal cost and impacts.

In order to provide packet producers with the capability to adjust the rate as well as

the contents of telemetry packets, a concept known as "zero-length packets" (ZLPs) was introduced. ZLPs are packets produced when an instrument has not collected enough information to provide a packet at the scheduled packet pickup. ZLPs are identified by setting the length field in the packet header to an invalid (too small) value, but the actual size of the packet is the size defined for that producer in the current STM.ZLPs arc collected by the CDS, but arc not stored or downlinked. This allows instruments to usc only as much of their packet collection rate allocation in a given STM as needed. The packet collection rate allocation is usually set at the maximum data rate the instrument is allowed to produce in that STM, so the instruments can vary the effective output data rate without affecting the spacecraft telemetry mode. T "his allows instruments with variable-ratio compression algorithms to use less downlink bandwidth and avoid storing packets which contain no data on the SSR. This capability, which gives the packet producers limited control over the volume of data produced, is a significant step toward allowing demand-driven telemetry.

### 3.0 Resulting Capabilities

As a result of the implementation of the Cassini telemetry system, many of the capabilities envisioned by the CCSDS have been made possible within the mission and architectural constraints. Additionally, there have been some resulting benefits which increase the quality of the spacecraft telemetry.

While the Cassini spacecraft telemetry system began as a very limited implementation of the CCSDS standards and philosophies, the additional capabilities which were incorporated show an evolution toward the CCSDS design concepts. The STMs used by Cassini limit the number of telemetry producers at a given time and also dictate the amount of data each producer is allowed to produce. However, the zero-length packet concept allows a telemetry producer to vary the actual data volume within the limit set by the STM. In addition, changing the packet type or the use of mini-packets gives telemetry producers a great deal of control over

the telemetry produced. When these concepts are used effectively by the telemetry producer, the very efficient use of the limited telemetry resources can be accomplished.

in addition, the mini-packet concept provided a secondary benefit to the end-users of the spacecraft telemetry - greater measurement intercorrelation. One of the results of the large size of the Cassini telemetry packets is coarse indication of measurement sampling time. Even at high downlink rates, several seconds may pass between the collection of successive packets from an engineering subsystem. Because measurements arc time-tagged with the packet time, it is difficult to determine exactly when a measurement changes values or when measurements within a packet were sampled with respect to onc another. However, minipackets can be used to sample measurements faster than once per packet. This allows for much better resolution of important measurements and better intercorrelation between measurements because the mini-packet time-tag applies to a smaller number of measurements.

#### 4.0 Conclusions

'l<hc CCSDS telemetry standards can be successfully applied to deep-space planetary missions, even in an environment of rigorous testing, design conservatism and biased culture. The Cassinitelemetry system is an example of this, as it incorporates the CCSDS protocols and even with severe architectural constraints, is a flexible telemetry system not unlike the systems envisioned by the CCSDS designers. This was accomplished by developing solutions which allowed variability in the spacecraft telemetry while remaining within the architectural tenets of the design and the constraints of the mission. The Cassini design solutions, and the solutions used on the Mars Observer spacecraft, reflect an evolution in IPL spacecraft telemetry systems away from the traditional time-division-multiplexed systems toward systems which completely embrace the CCSDS standards and philosophy.

### 5.0 Acknowledgments

The work described in this paper was carried out at the Jet l'ropulsion Laboratory, California Institute cjf]'ethnology, under a contract with the National Aeronautics and Space Administration. The implementation of the Cassini spacecraft telemetry system is based upon carly work in the spacecraft information system done by Sandy Kramer.

### 6.0 References

- 1 "Summary of Concept and Rationale", Report CCSDS100.0-G-1, Issuc 1, Green Book, Consultative Committee for Space Data Systems, December 1987.
- J. C. Day, "Telemetry Formats and Measurements", JPLPD 699-205-3-281, May 1994(JPL Internal I Document)